

# VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE : 8 May, 1998

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

Please read Appendix C of Announcement of Opportunity for details on how to fill in this Cover Sheet.

(1) Date prepared : 15-Apr-1998

(2) Proposal title : Polarization observations of 1055+018 at 1.6 and 5 GHz

(3)	INVESTIGATORS	INSTITUTION
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(5) Proposal Abstract :

We propose single-epoch polarization observations of the quasar 1055+018 at 1.6 and 5 GHz. Global VLBI observations at 1.7 GHz indicate that the millisecond-scale structure is roughly perpendicular to the arcsecond-scale structure mapped by VLA observations (they are offset by 60–70 degrees). In addition, the source is extremely compact and substantially polarized. The proposed HALCA measurements would make it possible to study the behavior of the position angle of the jet structure as it extends from the VLBI core to arcsecond scales. The availability of linear polarization measurements will provide additional information about the magnetic field geometry along the apparently curved jet.

(6) Proposal Category (indicate all that apply):

Object type:

☒ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other :

Observation type:

☐ Continuum, ☐ Spectral Line, ☒ Polarization, ☐ Time-critical, ☐ Other :

### (7) Number of proposed experiments

An ‘experiment’ is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a ‘monitoring experiment’ – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 2

### (8) Details of proposed experiments

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name ( <i>Jhhmm±ddmm</i> )	J1058+0133	J1058+0133		
Alternative name	1055+018	1055+018		
RA(J2000) (hh mm ss.ssss)	10 58 29.6100	10 58 29.6100		
Dec(J2000) (dd mm ss.ssss)	1 33 58.8000	1 33 58.8000		
Observing frequency band (GHz)	1.6	5		
<i>Continuum observations:</i>				
Standard VSOP freq. channels?	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Channel A range (MHz)				
Channel B range (MHz)				
<i>Spectral line observations:</i>				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel	32	32		
Correlator averaging time (sec)	2	2		
No. of correlating passes (if >1)				
Total flux density (Jy)	3.0	3.0		
Correlated flux (mJy)	1327	1327		
<i>Ground Radio Telescopes:</i>				
Suggested array given at Item (10)?	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>GRT observing mode:</i>				
128Mbps LCP (standard)	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
128Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
256Mbps LCP/RCP	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Preferred correlator:</i>				
No preference	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Mitaka	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Penticton	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Socorro	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Monitoring programs:</i>				
Number of observations				
Mean interval (days)				
Related AO1 proposal code(s)				

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

- ☒ 2 channel x 16 MHz, 2-bit (Standard mode),  
☐ Other:

Phase calibration tones:

- ☒ On (Standard continuum mode),  
☐ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (10) below)

(10) Additional notes to the scheduler :

VLBA + VLA (or Bonn) for both experiments  
possible observation date for 1055+018 May 1999

(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of  $(u,v)$  plots per source may optionally be included.  
(Refer to the VSOP Announcement of Opportunity for detailed instructions.)  
Preprints and reprints will not be forwarded to the Scientific Review Committee.

Send two paper copies of the complete proposal to:

VSOP Observing Proposals  
VSOP Science Operations Group  
Institute of Space and Astronautical Science  
3-1-1 Yoshinodai, Sagami-hara  
Kanagawa 229-8510 JAPAN

In addition, e-mail the completed L<sup>A</sup>T<sub>E</sub>X file to [submit@vsop.isas.ac.jp](mailto:submit@vsop.isas.ac.jp)

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

**Proposals must be received at ISAS by 8 May 1998**