## VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE	:	8	May,	1998
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SEND TO: VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

Please read Appendix C of Announcement of Opportunity for details on how to fill in this Cover Sheet.

(1) Date prepared: 28 April 1998

(2) Proposal title: Space-VLBI observations of the quasar 3C395

(3)	INVESTIGATORS	INSTITUTION
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(4) Principal Investigator (or contact person) details...

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(5) Proposal Abstract:

We propose VSOP observations of the quasar 3C395 simultaneously at 1.6 and 5 GHz. High resolution observations at 1.6 CHz will trace in detail the trajectory of the jet in its way "through"

stationary component B, at 15 mas from the core of 3C395 and with a rather steep spectru At 5 GHz, we will be able to study the details of the compact core-jet region of 3C395, a compare with recent VSOP observations at this frequency and VLBA observations at 15 GI The activity in this region is responsible of the total flux density variability observed at seve frequencies.	im. and Hz.
(C) Duran and Clata many (in directs all that any lay).	
(6) Proposal Category (indicate all that apply):  Object type:  ✓ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other:  Observation type:  ✓ Continuum, ☐ Spectral Line, ☐ Polarization, ☐ Time-critical, ☐ Other:	

## (7) Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 1

## (8) Details of proposed experiments

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name $(Jhhmm \pm ddmm)$	J1902+3159			
Alternative name	3C395			
RA(J2000) (hh mm ss.ssss)	19 02 55.93912			
Dec(J2000) (dd mm ss.ssss)	+31 59 41.7014			
Observing frequency band (GHz)	1.6/5			
Continuum observations:				
Standard VSOP freq. channels?				
Channel A range (MHz)				
Channel B range (MHz)				
Spectral line observations:				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes (if $>1$ )				
Total flux density (Jy)	2.5/1.6			
Correlated flux (mJy)	600/400			
Ground Radio Telescopes:				
Suggested array given at Item (10)?				
GRT observing mode:				
128Mbps LCP (standard)				
128Mbps LCP/RCP				
256Mbps LCP/RCP				
Preferred correlator:				
No preference				
Mitaka				
Penticton				
Socorro				
Monitoring programs:				
Number of observations				
Mean interval (days)				
Related AO1 proposal code(s)	v012			

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):
(10) Additional notes to the scheduler:
Suggested array: VLBA, MC, RO Suggested dates: from 25 September 1999 to 15 October 1999
(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of (u,v) plots per source may optionally be included. (Refer to the VSOP Announcement of Opportunity for detailed instructions.) Preprints and reprints will not be forwarded to the Scientific Review Committee.
Send two paper copies of the complete proposal to:  VSOP Observing Proposals  VSOP Science Operations Group

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

Proposals must be received at ISAS by 8 May 1998

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In addition, e-mail the completed LATEX file to submit@vsop.isas.ac.jp