VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE : 8 May, 1998 SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : 4th May 1998

(2) Proposal title : Cores of Lobe-dominated Quasars with VSOP - Epoch 2

(3)	INVESTIGATORS	INSTITUTION
P.I.	R.W.Porcas	MPIfR, Bonn, GERMANY
co-I.	J.A.Zensus	MPIfR, Bonn, GERMANY
co-I.	D.H.Hough	Trinity University, San Antonio, TX, USA
co-I.		

(4) Principal Investigator (or contact person) details...

Name : Richard Porcas	$\operatorname{Address}$: MPIfR
E-mail · porcas@mpifr-bonn mpg de	: Auf dem Huegel 69,
Fax \cdot 49 228 525229	: D 53121 Bonn
Phone : 49 228 525288	: GERMANY
	:

(5) Proposal Abstract :

The study of the parsec-scale morphology and kinematics of quasars is central to the understanding of extragalactic radio source jet physics, and the classification of various types of objects. In the "unified schemes" the properties of sources should depend crucially upon the angle of a relativistically moving plasma with respect to the line of sight. Core-dominated sources should show properties corresponding to an axis nearly pointing at the observer - superluminal motion, exaggeration of small bends, and strong and variable cores. In contrast, in lobe- dominated quasars these properties should be weak or absent. We wish to make SECOND EPOCH VSOP OBSERVATIONS of our small sample of LOBE-DOMINATED QUASARS, to compare the submas structures and component velocities with those of the stronger core-dominated quasars. Our sources are at high declinations, and are thus ideal for efficient use of VSOP-GRT time.

(6) Proposal Category (indicate all that apply):					
Object type:					
\checkmark AGN, \square Maser, \square Stellar, \square Pulsar, \square Other :					
Observation type:					
\checkmark Continuum, \square Spectral Line, \square Polarization, \square Time-critical, \square Other :					

(7) Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 3

(8) Details of proposed experim

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name $(Jhhmm \pm ddmm)$	J0728+6748	J1139+6547	J0825+6157	-
Alternative name	3C179	3C263		
RA(J2000) (hh mm ss.ssss)	07 28 11.566	11 39 57.043	$08 \ 25 \ 38.643$	
Dec(J2000) (dd mm ss.ssss)	$+67 \ 48 \ 47.19$	+65 47 49.25	$+61\ 57\ 28.68$	
Observing frequency band (GHz)	5	5	5	
Continuum observations:				
Standard VSOP freq. channels?	∇	∇	∇	
Channel A range (MHz)				
Channel B range (MHz)				
Spectral line observations:				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes $(if > 1)$				
Total flux density (Jy)	0.6	0.2	0.6	
Correlated flux (mJy)	400	150	400	
Ground Radio Telescopes:				
Suggested array given at Item (10) ?	∇	∇	∇	
GRT observing mode:				
128Mbps LCP (standard)	∇	$\overline{\mathbf{V}}$	∇	
128Mbps LCP/RCP				
256 Mbps LCP/RCP				
Preferred correlator:				
No preference				
Mitaka				
Penticton				
Socorro				
Monitoring programs:				
Number of observations				
Mean interval (days)				
Related AO1 proposal code(s)	V054	V054	V054	

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

2 channel x 16 MHz, 2-bit (Standard mode),

Other:

Phase calibration tones:

On (Standard continuum mode),

Off (Standard spectral line mode)

(Include justification of any non-standard choice at (10) below)

(10) Additional notes to the scheduler :

Because of the weak nature of these sources, and the desire to have 15 GHz VLBA "snapshot" observations in the tracking station gaps, we strongly request that our array of choice be used. Experiment 1: VLBA + EB Experiment 2: VLBA + EB Experiment 3: VLBA + EB

(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of (u,v) plots per source may optionally be included. (Refer to the VSOP Announcement of Opportunity for detailed instructions.) Preprints and reprints will not be forwarded to the Scientific Review Committee.

Send two paper copies of the complete proposal to: VSOP Observing Proposals VSOP Science Operations Group Institute of Space and Astronautical Science 3-1-1 Yoshinodai, Sagamihara Kanagawa 229-8510 JAPAN In addition, e-mail the completed IATEX file to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

Proposals must be received at ISAS by 8 May 1998