VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE : 8 May, 1998 SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

Please read Appendix C of Announcement of Opportunity for details on how to fill in this Cover Sheet.

(1) Date prepared : 8-May-1998

(2) Proposal title : The parsec-scale kinematics of BL Lac objects PKS 1413+135

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(5) Proposal Abstract :

PKS 1413+135 has properties of BL Lac Objects and a parsec-scale counterjet. This duality makes this object to be unique. Although the jet motion is a key to investigate the nature of such unique source, the motion of any jet-component has not been accurately measured by VLBI up to now. We propose imaging observations with VSOP of this object. It is possible to first obtain the jet motion of PKS 1413+135 by VSOP. This study allows us to clarify the kinematics of parsec-scale radio jets in BL Lac objects.

(6) Proposal Category (indicate all that apply):						
Object type:						
$\overrightarrow{\nabla}$ AGN, \square Maser, \square Stellar, \square Pulsar, \square Other :						
Observation type:						
Continuum, Spectral Line, Polarization, Time-critical, Oth	ier :					

(7) Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 2

(8) Details of proposed experim

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name $(Jhhmm \pm ddmm)$	J1415 + 1320	J1415 + 1320		
Alternative name	PKS 1413+135	PKS 1413+135		
RA(J2000) (hh mm ss.ssss)	$14 \ 15 \ 58.82$	$14 \ 15 \ 58.82$		
Dec(J2000) (dd mm ss.ssss)	$+13 \ 20 \ 23.69$	$+13 \ 20 \ 23.69$		
Observing frequency band (GHz)	5	1.6		
Continuum observations:				
Standard VSOP freq. channels?		∇		
Channel A range (MHz)				
Channel B range (MHz)				
Spectral line observations:				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes $(if > 1)$				
Total flux density (Jy)	0.92	1.07		
Correlated flux (mJy)	547	243		
Ground Radio Telescopes:				
Suggested array given at Item (10) ?	∇	$\overline{\mathbf{A}}$		
GRT observing mode:				
128Mbps LCP (standard)		$\overline{\mathbf{A}}$		
128Mbps LCP/RCP				
256Mbps LCP/RCP				
Preferred correlator:				
No preference		$\overline{\nabla}$		
Mitaka				
Penticton				
Socorro				
Monitoring programs:				
Number of observations				
Mean interval (days)				
Related AO1 proposal code(s)				

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

✓ 2 channel x 16 MHz, 2-bit (Standard mode),
Other:

Phase calibration tones:

✓ On (Standard continuum mode),
✓ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (10) below)

(10) Additional notes to the scheduler :

A high dynamic range is required in this project. Experiment 1, Pref Array = VLBA, Kashima, Mopra, Shanghai Experiment 2, Pref Array = VLBA, VLA, Tidbinbilla, Kashima, Shanghai

(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of (u,v) plots per source may optionally be included. (Refer to the VSOP Announcement of Opportunity for detailed instructions.) Preprints and reprints will not be forwarded to the Scientific Review Committee.

Send two paper copies of the complete proposal to: VSOP Observing Proposals VSOP Science Operations Group Institute of Space and Astronautical Science 3-1-1 Yoshinodai, Sagamihara Kanagawa 229-8510 JAPAN In addition, e-mail the completed LATEX file to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

Proposals must be received at ISAS by 8 May 1998