

# VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE : 8 May, 1998

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : 8-May-1998

(2) Proposal title : VSOP Observations of AO 0235+164: extreme high-speed jet?

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(5) Proposal Abstract :

BL Lac object AO 0235+164 is a strong variable Active Galactic Nuclei (AGNs) from radio to gamma-ray energy band. A number of VLBI studies showed the source has a very compact ( $< 0.1$  mas) core and weak jet component with fast expansion speed. We have shown that the source possibly has Doppler factor  $\delta \geq 119$ , which is the largest known  $\delta$  to date. This large  $\delta$ , suggesting very high-speed jet, could be the cause of variability and other activities.

We propose VSOP monitoring observations of 0235+164 which allow us to detect the possible high-speed jet as very fast superluminal motion. We can test the jet acceleration and propagation mechanisms with such high-speed jet.

(6) Proposal Category (indicate all that apply):

Object type:

☒ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other :

Observation type:

☒ Continuum, ☐ Spectral Line, ☐ Polarization, ☐ Time-critical, ☐ Other :

**(7) Number of proposed experiments**

An ‘experiment’ is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a ‘monitoring experiment’ – requires only one column to be filled in. Suggested observing dates, especially for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 2

**(8) Details of proposed experiments**

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name ( <i>Jhhmm±ddmm</i> )	J0238+1636	J0238+1636		
Alternative name				
RA(J2000) (hh mm ss.ssss)	02 38 38.9301	02 38 38.9301		
Dec(J2000) (dd mm ss.ssss)	+16 36 59.2752	+16 36 59.2752		
Observing frequency band (GHz)	5	5		
<i>Continuum observations:</i>				
Standard VSOP freq. channels?	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Channel A range (MHz)				
Channel B range (MHz)				
<i>Spectral line observations:</i>				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes (if >1)				
Total flux density (Jy)	2	2		
Correlated flux (mJy)	1500	1500		
<i>Ground Radio Telescopes:</i>				
Suggested array given at Item (10)?	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>GRT observing mode:</i>				
128Mbps LCP (standard)	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
128Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
256Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Preferred correlator:</i>				
No preference	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Mitaka	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Penticton	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Socorro	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Monitoring programs:</i>				
Number of observations	4	2		
Mean interval (days)	14	28		
Related AO1 proposal code(s)				

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

- ☒ 2 channel x 16 MHz, 2-bit (Standard mode),  
☐ Other:

Phase calibration tones:

- ☒ On (Standard continuum mode),  
☐ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (10) below)

(10) Additional notes to the scheduler :

Preferred array = VLBA + some telescopes located apart from North America.

Preferred scheduling date:

Exp. 1 = Jan 1999

Exp. 2 = Aug 1999

(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of  $(u,v)$  plots per source may optionally be included.

(Refer to the VSOP Announcement of Opportunity for detailed instructions.)

Preprints and reprints will not be forwarded to the Scientific Review Committee.

Send two paper copies of the complete proposal to:

VSOP Observing Proposals

VSOP Science Operations Group

Institute of Space and Astronautical Science

3-1-1 Yoshinodai, Sagami-hara

Kanagawa 229-8510 JAPAN

In addition, e-mail the completed L<sup>A</sup>T<sub>E</sub>X file to [submit@vsop.isas.ac.jp](mailto:submit@vsop.isas.ac.jp)

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

**Proposals must be received at ISAS by 8 May 1998**