

# VSOP AO2 PROPOSAL COVER SHEETS

DEADLINE : 8 May, 1998

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : 8-May-1998

(2) Proposal title : VSOP Observations of 1611+343: Gamma-ray emission mechanism

(3)	INVESTIGATORS	INSTITUTION
P.I.	Kenta Fujisawa	ISAS, Japan
co-I.	H.Kobayashi, H.Hirabayashi	ISAS, Japan
co-I.	K. Wajima	Ibaraki Univ., Mito, Japan
co-I.	S. Kamenno, M. Inoue, S. Iguchi	NAO, Mitaka, Japan
co-I.		
co-I.		
co-I.		
co-I.		
co-I.		

(4) Principal Investigator (or contact person) details...

Name : Kenta Fujisawa	Address : ISAS
E-mail : kenta@vsop.isas.ac.jp	: 3-1-1, Yoshinodai, Sagamihara
Fax : +81-427-51-3972	: Kanagawa, 229-8510
Phone : +81-427-51-3911	: Japan
	:

(5) Proposal Abstract :

On the basis of the CGRO/EGRET observations, it has been discussed that a possible connection between the high energy gamma-ray activity and high-speed radio jet of AGN. However there seems no definite evidence of such beaming.

We propose to observe a gamma-ray quasar 1611+343 by VSOP, to study clear connection of radio jet and gamma-ray emission. According to the ground-based VLBI observations, this source possibly has a faster jet close to the central core, while the outer jet is relatively slow. Since the gamma-ray is emitted at or close to the core, the existence of faster jet in a small scale strongly supports the beamed gamma-ray hypothesis.

(6) Proposal Category (indicate all that apply):

Object type:

☒ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other :

Observation type:

☒ Continuum, ☐ Spectral Line, ☐ Polarization, ☐ Time-critical, ☐ Other :

**(7) Number of proposed experiments**

An ‘experiment’ is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a ‘monitoring experiment’ – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (10).

The number of experiments in this proposal is: 1

**(8) Details of proposed experiments**

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name ( <i>Jhhmm±ddmm</i> )	J1613+3412			
Alternative name	1611+343			
RA(J2000) (hh mm ss.ssss)	16 13 41.0642			
Dec(J2000) (dd mm ss.ssss)	+34 12 47.9089			
Observing frequency band (GHz)	5			
<i>Continuum observations:</i>				
Standard VSOP freq. channels?	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Channel A range (MHz)				
Channel B range (MHz)				
<i>Spectral line observations:</i>				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes (if >1)				
Total flux density (Jy)	4			
Correlated flux (mJy)	2000			
<i>Ground Radio Telescopes:</i>				
Suggested array given at Item (10)?	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>GRT observing mode:</i>				
128Mbps LCP (standard)	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
128Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
256Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Preferred correlator:</i>				
No preference	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Mitaka	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Penticton	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Socorro	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Monitoring programs:</i>				
Number of observations	2			
Mean interval (days)	60			
Related AO1 proposal code(s)				

(9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

- ☒ 2 channel x 16 MHz, 2-bit (Standard mode),  
☐ Other:

Phase calibration tones:

- ☒ On (Standard continuum mode),  
☐ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (10) below)

(10) Additional notes to the scheduler :

Preferred array = VLBA + some telescopes located apart from North America.

Preferred scheduling date = Jun – Aug, 1999

(11) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Up to one page of  $(u,v)$  plots per source may optionally be included.

(Refer to the VSOP Announcement of Opportunity for detailed instructions.)

Preprints and reprints will not be forwarded to the Scientific Review Committee.

Send two paper copies of the complete proposal to:

VSOP Observing Proposals

VSOP Science Operations Group

Institute of Space and Astronautical Science

3-1-1 Yoshinodai, Sagami-hara

Kanagawa 229-8510 JAPAN

In addition, e-mail the completed L<sup>A</sup>T<sub>E</sub>X file to [submit@vsop.isas.ac.jp](mailto:submit@vsop.isas.ac.jp)

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

**Proposals must be received at ISAS by 8 May 1998**