VSOP AO3 PROPOSAL COVER SHEETS

DEADLINE : 1 October, 1999

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : 1 October, 1999

(2) Proposal title : Wiggling structure of Double Sided Jets in 0458+138

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(5) Proposal Abstract :

We propose observations of 0458+138. This is a unique source having symmetrical jets at the opposite sides of its core component. Those jets are continuously ejected from the core. Our main purpose of this proposal is to detect a "wiggle-like" structure in each of jets. In this study, we investigate mechanisms of the accerelation and collimation of VLBI jets. If the magnetic field is dominant to produce wiggling jets, the wiggle pattern will show mirror-symmetry in respect with the core. On the other hand, if the precession of an accretion disk is dominant, the wiggle will show point-symmetry. The curving of the jets are slightly seen in the opposite sided jets at 8.5GHz, so that VSOP observations will reveal the wiggle-like structure of 0458+138.

(6) Proposal Category (indicate all that apply):
Object type:
\checkmark AGN, \square Maser, \square Stellar, \square Pulsar, \square Other :
Observation type:
\checkmark Continuum, \square Spectral Line, \square Polarization, \square Time-critical, \square Other :

(7) Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (11).

The number of experiments in this proposal is:1

(8) Details of proposed experiments

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name $(Jhhmm \pm ddmm)$	J0501+1356	J0501+1356		
Alternative name				
RA(J2000) (hh mm ss.ssss)	$05 \ 01 \ 45.2708$	$05 \ 01 \ 45.2708$		
Dec(J2000) (dd mm ss.ssss)	+13 56 07.221	+13 56 07.221		
Observing frequency band (GHz)	1.6	5		
Continuum observations:				
Standard VSOP freq. channels?	∇	∇		
Channel A range (MHz)				
Channel B range (MHz)				
Spectral line observations:				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel	64	64		
Correlator averaging time (sec)	2	2		
No. of correlating passes $(if > 1)$				
Total flux density (Jy)	0.6	0.6		
Correlated flux (mJy)	270	270		
Ground Radio Telescopes:				
Suggested array given at Item (11) ?	∇	∇		
GRT observing mode:				
128Mbps LCP (standard)	∇	$\overline{\mathbf{V}}$		
128Mbps LCP/RCP				
256 Mbps LCP/RCP				
Preferred correlator:				
No preference	∇	∇		
Mitaka				
Penticton				
Socorro				
Monitoring programs:				
Number of observations				
Mean interval (days)				
Related VSOP proposal code(s)				

- (9) VSOP spacecraft observing mode (see Section 3 and Table 5 of the VSOP Proposer's Guide):

 ✓ 2 channel x 16 MHz, 2-bit (Standard mode),
 Other:

 Phase calibration tones:
 ✓ On (Standard continuum mode),
 - Off (Standard spectral line mode),

(Include justification of any non-standard choice at (11) below)

- (11) Additional notes to the scheduler :

VLBA as a ground array is expected. In addition, one large telescope is expected.

(12) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Refer to the VSOP Announcement of Opportunity for detailed instructions. Preprints and reprints will not be forwarded to the Scientific Review Committee.

EITHER e-mail the completed $L^{AT}EX$ file to submit@vsop.isas.ac.jp and send two paper copies of the complete proposal to:

VSOP Observing Proposals VSOP Science Operations Group Institute of Space and Astronautical Science 3-1-1 Yoshinodai, Sagamihara Kanagawa 229-8510 JAPAN

 \mathbf{OR} e-mail the completed LATEX Cover Sheets file and, in a separate e-mail, the postscript file of the scientific and technical justification, to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

Proposals must be received at ISAS by 1 October 1999