# VSOP AO4 PROPOSAL COVER SHEETS

DEADLINE : 2 October, 2000
SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

**(1)** Date prepared : September 29, 2000

**(2)** Proposal title : The Second-Epoch Observations of 3C 380

| **(3)** | INVESTIGATORS | INSTITUTION |
|---|---|---|
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**(5)** Proposal Abstract :


We propose the second-epoch observations of the quasar 3C 380 at 1.7 and 5 GHz. In the first-epoch observations on July 1998, we confirmed superluminal motions of the jet by comparing with ground VLBI. We found that the apparent velocities of components F and A were significantly different from each other in terms of the speed and position angle, while each component keeps each constant speed and straight motion for more than 15 years. We also found shell-like enhancements of the spectral indices at these components. The result can be evidence of ballistic adeabatique expansion of plasma blobs. The second-epoch VSOP observations, separated by 2.5 years, will measure the jet motions as accurate as the ground 15-year monitor, and allow us to investigate how straight the motion is (or acceleration, if any).


**(6)** Proposal Category (indicate all that apply):

Object type:
☑ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other :

Observation type:
☑ Continuum, ☐ Spectral Line, ☐ Polarization, ☐ Time critical,
☐ Phase-reference, ☐ Other :

**(7)** Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (11).

The number of experiments in this proposal is:2

**(8)** Details of proposed experiments

| | Experiment 1 | Experiment 2 | Experiment 3 | Experiment 4 |
|---|---|---|---|---|
| Source name (J*hhmm*±*ddmm*) | J1829+4843 | J1829+4843 | | |
| Alternative name | 3C 380 | 3C 380 | | |
| RA(J2000) (hh mm ss.ssss) | 18 29 31.804 | 18 29 31.804 | | |
| Dec(J2000) (dd mm ss.sss) | +48 44 46.62 | +48 44 46.62 | | |
| Observing frequency band (GHz) | 1.6 | 5 | | |
| *Continuum observations:* | | | | |
| Standard VSOP freq. channels? | ☑ | ☑ | ☐ | ☐ |
| Channel A range (MHz) | | | | |
| Channel B range (MHz) | | | | |
| *Spectral line observations:* | | | | |
| Ch.A spectral line rest freq. (MHz) | | | | |
| Ch.A LSR velocity (km/s) | | | | |
| Ch.B spectral line rest freq. (MHz) | | | | |
| Ch.B LSR velocity (km/s) | | | | |
| FWHM of field of view required (mas) | | | | |
| Min. spectral channels per IF channel | 128 | 128 | | |
| Correlator averaging time (sec) | | | | |
| No. of correlating passes (if >1) | | | | |
| Total flux density (Jy) | 9.5 | 7.5 | | |
| Correlated flux (mJy) | 100 – 800 | 100 – 800 | | |
| *Ground Radio Telescopes:* | | | | |
| Suggested array given at Item (11)? | ☑ | ☑ | ☐ | ☐ |
| *GRT observing mode:* | | | | |
| 128Mbps LCP (standard) | ☑ | ☑ | ☐ | ☐ |
| 128Mbps LCP/RCP | ☐ | ☐ | ☐ | ☐ |
| 256Mbps LCP/RCP | ☐ | ☐ | ☐ | ☐ |
| *Preferred correlator:* | | | | |
| No preference | ☑ | ☑ | ☐ | ☐ |
| Mitaka | ☐ | ☐ | ☐ | ☐ |
| Penticton | ☐ | ☐ | ☐ | ☐ |
| Socorro | ☐ | ☐ | ☐ | ☐ |
| *Monitoring programs*: | | | | |
| Number of observations | | | | |
| Mean interval (days) | | | | |
| Related VSOP proposal code(s) | V125a | V125b | | |

**(9)** VSOP spacecraft observing mode (see Section 3 and Table 2 of the VSOP Proposer's Guide):

☑ 2 channel x 16 MHz, 2-bit (Standard mode),
☐ Other:

Phase calibration tones:

☑ On (Standard continuum mode),
☐ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (11) below)

**(10)** Assistance with preparation of ground telescope schedule files:

☐ VSOG assistance requested, ☐ Consultation desired, ☑ No assistance required

**(11)** Additional notes to the scheduler :

Since 3C 380 showes an extended structure, visibility amplitudes of the space baselines rapidly vary in the range of 0.1 – 0.8 Jy. We request attendance of at least one large telescope such as the phased VLA and Effelsberg, to salvage fringes at the minimums of visibilities.

**(12)** Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Refer to the VSOP Announcement of Opportunity for detailed instructions. Preprints and reprints will not be forwarded to the Scientific Review Committee.

**EITHER** e-mail the completed LaTeX file to submit@vsop.isas.ac.jp and send two paper copies of the complete proposal to:

VSOP Observing Proposals
VSOP Science Operations Group
Institute of Space and Astronautical Science
3-1-1 Yoshinodai, Sagamihara
Kanagawa 229-8510 JAPAN

**OR** e-mail the completed LaTeX Cover Sheets file and, in a separate e-mail, the postscript file of the scientific and technical justification, to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

**Proposals must be received at ISAS by 2 October 2000**