## **VSOP AO5 PROPOSAL COVER SHEETS**

DEADLINE : 1 February, 2001

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : January 30, 2001

(2) Proposal title : Pair plasma dominance in the pc-scale jets of B1150+812 and B1213+350?

(3)	INVESTIGATORS	INSTITUTION
P.I.	Kouichi Hirotani	National Astronomical Observatory, Tokyo, Japan
co-I.	M.A. Pérez-Torres	Istituto di Radioastronomia, Bologna, Italy
co-I.	S. Kameno	National Astronomical Observatory, Tokyo, Japan
co-I.	J.M. Marcaide	Universidad de Valencia, Spain
co-I.		

(4) Principal Investigator (or contact person) details...

Name : Kouichi Hirotani Address E-mail : hirotani@hotaka.mtk.nao.ac.jp Fax : +81-422-34-3814 Phone : +81-422-34-3941

Address : National Astronomical

- : Observatory,
- : Osawa 2-21-1, Mitaka,
- : Tokyo 181-8588,
- : Japan

(5) Proposal Abstract :

We propose VSOP observations of the quasars B1150+812 and B1213+350 at 18cm, aimed at discriminating whether their parsec-scale jets are dominated by a "normal plasma" (protrons and relativistic electrons), or by a "pair plasma" (relativistic electrons and positrons). For this study, we need the synchrotron self-absorption spectra for the individual jet components of each quasar. Therefore, we also propose quasi-simultaneous observations with the VLBA (at 13, 6, 3.6, and 2 cm).

While the multifrequency VLBA observations will allow us to accurately constrain the spectra of the sources, we also need to modelfit and decompose the components from 18 cm to 2 cm. Hence, VSOP observations at the longest wavelength, 18 cm, are mandatory.

(6) Proposal Category (indicate all that apply):					
Object type:					
$\overrightarrow{V}$ AGN, $\square$ Maser, $\square$ Stellar, $\square$ Pulsar, $\square$ Other :					
Observation type:					
$\bigtriangledown$ Continuum, $\square$ Spectral Line, $\square$ Polarization, $\square$ Time critical,					
Phase-reference, Other:					

## (7) Number of proposed experiments

An 'experiment' is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a 'monitoring experiment' – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (11).

The number of experiments in this proposal is: 2

## (8) Details of proposed experiments

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name $(Jhhmm \pm ddmm)$	J1153+8058	J1215+3548		
Alternative name				
RA(J2000) (hh mm ss.ssss)	11 53 12.4993	$12 \ 15 \ 55.6010$		
Dec(J2000) (dd mm ss.sss)	$+80\ 58\ 29.154$	+34 48 15.220		
Observing frequency band (GHz)	1.6	1.6		
Continuum observations:				
Standard VSOP freq. channels?	$\nabla$	$\nabla$		
Channel A range (MHz)				
Channel B range (MHz)				
Spectral line observations:				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity $(km/s)$				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes $(if > 1)$				
Total flux density (Jy)	1.4	1.3		
Correlated flux (mJy)	800	800		
Ground Radio Telescopes:				
Suggested array given at Item $(11)$ ?				
GRT observing mode:				
128Mbps LCP (standard)	$\nabla$	$\nabla$		
128Mbps LCP/RCP				
256 Mbps LCP/RCP				
Preferred correlator:				
No preference	$\overline{\mathbf{V}}$	$\overline{\mathbf{A}}$		
Mitaka				
Penticton				
Socorro				
Monitoring programs:				
Number of observations				
Mean interval (days)				
Related VSOP proposal code(s)				

- - $\nabla$  On (Standard continuum mode),
  - Off (Standard spectral line mode)

(Include justification of any non-standard choice at (11) below)

- (10) Assistance with preparation of ground telescope schedule files:
  ✓ VSOG assistance requested, Consultation desired, No assistance required
- (11) Additional notes to the scheduler :

(12) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Refer to the VSOP Announcement of Opportunity for detailed instructions. Preprints and reprints will not be forwarded to the Scientific Review Committee.

**EITHER** e-mail the completed  $IAT_EX$  file to submit@vsop.isas.ac.jp and send two paper copies of the complete proposal to:

VSOP Observing Proposals VSOP Science Operations Group Institute of Space and Astronautical Science 3-1-1 Yoshinodai, Sagamihara Kanagawa 229-8510 JAPAN

 $\mathbf{OR}$  e-mail the completed  $\mathbb{IAT}_{EX}$  Cover Sheets file and, in a separate e-mail, the postscript file of the scientific and technical justification, to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

Proposals must be received at ISAS by 1 February 2001