

# VSOP A05 PROPOSAL COVER SHEETS

DEADLINE : 1 February 2001

SEND TO : VSOG, ISAS, 3-1-1 Yoshinodai, Sagamihara, Kanagawa 229-8510, JAPAN

(1) Date prepared : 2000 February 1

(2) Proposal title : Monitoring PKS 1413+135 for Superluminal Motion and Spectral Evolution

(3)	INVESTIGATORS	INSTITUTION
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(5) Proposal Abstract :

We request three epochs of VLBA and HALCA observations of PKS1413+135 at 5 GHz, as well as contemporaneous VLBA-only observations at 15 GHz. Our goal is to confirm superluminal motion of the inner components of PKS1413+135 and detect the spectral change of these components. PKS1413+135 is one of the brightest red BL Lacs, and is an extremely compact radio source: total size  $\sim 400$  pc. VLBA and HALCA observations have revealed evolution on *both* sides of the nucleus, and suggest that there is superluminal motion of the inner most components of the source, but that outer components are stationary or slowly moving.

The super luminal motion is difficult to reconcile with unified schemes, unless the jet is strongly bent within a few milli-arcseconds of the core.

**(6)** Proposal Category (indicate all that apply):

Object type:

☒ AGN, ☐ Maser, ☐ Stellar, ☐ Pulsar, ☐ Other :

Observation type:

☒ Continuum, ☐ Spectral Line, ☐ Polarization, ☐ Time critical,☐ Phase-reference, ☐ Other :

**(7) Number of proposed experiments**

An ‘experiment’ is one or more observations of one source at a fixed HALCA set-up. A request to observe the same source at 1.6 GHz and separately at 5 GHz requires two columns to be filled in in item (8) below. A request to observe the same source with HALCA simultaneously observing at 1.6 GHz and 5 GHz requires one column to be filled in. Multi-epoch observations of the same source at the same frequency – a ‘monitoring experiment’ – requires only one column to be filled in. Suggested observing dates, especially for for time-critical and monitoring experiments, should be specified in item (11).

The number of experiments in this proposal is: 1

**(8) Details of proposed experiments**

	Experiment 1	Experiment 2	Experiment 3	Experiment 4
Source name ( <i>Jhhmm±ddmm</i> )	J1415+1320			
Alternative name				
RA(J2000) (hh mm ss.ssss)	14 15 58.817			
Dec(J2000) (dd mm ss.sss)	+13 20 23.713			
Observing frequency band (GHz)	5			
<i>Continuum observations:</i>				
Standard VSOP freq. channels?	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Channel A range (MHz)	16			
Channel B range (MHz)	16			
<i>Spectral line observations:</i>				
Ch.A spectral line rest freq. (MHz)				
Ch.A LSR velocity (km/s)				
Ch.B spectral line rest freq. (MHz)				
Ch.B LSR velocity (km/s)				
FWHM of field of view required (mas)				
Min. spectral channels per IF channel				
Correlator averaging time (sec)				
No. of correlating passes (if >1)				
Total flux density (Jy)	1			
Correlated flux (mJy)	1000			
<i>Ground Radio Telescopes:</i>				
Suggested array given at Item (11)?	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>GRT observing mode:</i>				
128Mbps LCP (standard)	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
128Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
256Mbps LCP/RCP	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Preferred correlator:</i>				
No preference	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Mitaka	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Penticton	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Socorro	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
<i>Monitoring programs:</i>				
Number of observations	3			
Mean interval (days)	15			
Related VSOP proposal code(s)	W083			

(9) VSOP spacecraft observing mode (see Section 3 and Table 2 of the VSOP Proposer's Guide):

- ☒ 2 channel x 16 MHz, 2-bit (Standard mode),  
☐ Other:

Phase calibration tones:

- ☒ On (Standard continuum mode),  
☐ Off (Standard spectral line mode)

(Include justification of any non-standard choice at (11) below)

(10) Assistance with preparation of ground telescope schedule files:

- ☐ VSOG assistance requested, ☒ Consultation desired, ☐ No assistance required

(11) Additional notes to the scheduler :

(12) Attach a scientific and technical justification, not in excess of 2 pages of text and 2 pages of figures. Refer to the VSOP Announcement of Opportunity for detailed instructions. Preprints and reprints will not be forwarded to the Scientific Review Committee.

**EITHER** e-mail the completed L<sup>A</sup>T<sub>E</sub>X file to submit@vsop.isas.ac.jp and send two paper copies of the complete proposal to:

VSOP Observing Proposals  
VSOP Science Operations Group  
Institute of Space and Astronautical Science  
3-1-1 Yoshinodai, Sagamihara  
Kanagawa 229-8510 JAPAN

**OR** e-mail the completed L<sup>A</sup>T<sub>E</sub>X Cover Sheets file and, in a separate e-mail, the postscript file of the scientific and technical justification, to submit@vsop.isas.ac.jp

Information from the Cover Sheets of scheduled proposals will be made available from the VSOP WWW site.

**Proposals must be received at ISAS by 2 October 2000**